

$\mathcal{O}(\alpha_s^3)$ calculations for the inclusive determination of $|V_{cb}|$

Zeuthen 2021

Kay Schönwald | November 4, 2021

TTP KARLSRUHE

[based on: Fael, Schönwald, Steinhauser PRL 125 (2020); JHEP 10 (2020); PRD 103 (2021); PRD 104 (2021)]





TRR 257 - Particle Physics Phenomenologyafter the Higgs Discovery

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Motivation

- b → cℓν is an important ingredient in the inclusive determination of |V_{cb}|:
 - Currently there is a tension between inclusive and exclusive determination of |V_{cb}|.
 - Errors are mostly theory dominated.
 - Precise measurements of the CKM matrix elements |V_{ib}| are among main goals of Belle II and LHCb.
 - The semi-leptonic decay rate is an important ingredient in the global fit for the inclusive determination.
 - The global fits are performed in the kinetic scheme.
- $\mu \rightarrow e \nu \nu$ is the most precise way to determine G_F .



Motivation

Introduction

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The Heavy Quark Expansion (HQE):

$$\mathrm{d}\Gamma = \mathrm{d}\Gamma_0 + \mathrm{d}\Gamma_{\mu_{\pi}}\frac{\mu_{\pi}^2}{m_b^2} + \mathrm{d}\Gamma_{\mu_G}\frac{\mu_G^2}{m_b^2} + \mathrm{d}\Gamma_{\rho_D}\frac{\rho_D^3}{m_b^3} + \mathrm{d}\Gamma_{\rho_{LS}}\frac{\rho_{LS}^3}{m_b^3} + \dots$$

- $d\Gamma_i$ are computed in perturbative QCD
- dependece on non-perturbative HQE parameters: $\mu_{\pi}, \mu_{G}, \rho_{D}, \rho_{LS}, \ldots$
- Perturbative corrections to $b \rightarrow c\ell\nu$ exhibit a bad convergence in the on-shell and the MS scheme for the heavy quark masses.
- Better knowledge of scheme conversion can further constrain the global fit.

The Kinetic Mass



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	tree	α_{s}	α_{s}^{2}	$lpha_{s}^{3}$		
1	\checkmark	\checkmark	\checkmark	this talk	[Jezabek, Kühn, NPB 314 (1989); Gambino et al., NPB 719 (2005)] [Melnikov, PLB 666 (2008); Pak, Czarnecki, PRD (2008)] [Fael, KS, Steinhauser, PRD 104 (2021)]	
1/m ² _b	\checkmark	\checkmark	ļ		[Alberti, Gambino, Nandi, JHEP 1401 (2014)] [Mannel, Pivovarov, Rosenthal, PRD 92 (2015)] [Becher, Boos, Lunghi, JHEP 0712 (2007)]	
$1/m_{b}^{3}$	\checkmark	\checkmark			[Mannel, Pivovarov, PRD 100 (2019)]	
1/m _b ^{4,5}	\checkmark				[Dassinger, Mannel, Turczyk, JHEP 0703 (2007); JHEP 1011 (2010)] [Fael, Mannel, Vos, JHEP 02 (2019); JHEP 12 (2019)]	
$\overline{m}_b - m_b^{ m kin}$		✓	\checkmark	this talk	[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)] [Fael, Steinhauser, KS, PRL 125 (2020); PRD 103 (2021)]	
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The Kinetic Heavy Quark Mass

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The kinetic heavy quark mass



- The kinetic mass scheme is tailored for $b \rightarrow c$ transitions.
- It is based on the heavy guark hadron mass relation in HQET:

Definition

$$M_B = m_b - \overline{\Lambda}(\mu) - rac{\mu_\pi^2(\mu)}{2m_b} + \dots$$

[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)]

• Key mass scale is m_b , not M_B :

$$\Gamma_{
m sl}\simeq rac{G_F^2|V_{cb}|^2}{192\pi^3}(M_B-\overline{\Lambda})^5$$

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The kinetic heavy quark mass



- The kinetic mass scheme is tailored for $b \rightarrow c$ transitions.
- It is based on the heavy quark hadron mass relation in HQET:

Definition

$$m^{ ext{kin}} = m^{ ext{OS}} - \overline{\Lambda}(\mu) ig|_{ ext{pert}} - rac{\mu_\pi^2(\mu) ig|_{ ext{pert}}}{2m^{ ext{kin}}} + \dots$$

[Bigi et al., PRD 56 (1997); Czarnecki, Melnikov, Uraltsev, PRL 80 (1998)]

• Key mass scale is *m*_b, not *M*_B:

$$\Gamma_{
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The Small Velocity Sum Rules





- $W(\omega, \vec{v}) = 2 \operatorname{Im} \left[\frac{i}{2m} \int d^4 x \, e^{-iqx} \langle Q | TJ(x)J(0) | Q \rangle \right]$ (*J* is an arbitrary current)
- \vec{v} : velocity of the heavy quark
- ω : excitation energy of the heavy quark

$$\overline{\Lambda}(\mu)\big|_{\text{pert}} = \lim_{\vec{v}\to 0} \lim_{m\to\infty} \frac{2}{\vec{v}^2} \frac{\int\limits_{0}^{\mu} d\omega \,\omega \, W(\omega, \vec{v})}{\int\limits_{0}^{\mu} d\omega \, W(\omega, \vec{v})}, \qquad \mu_{\pi}^2(\mu)$$

$$\mu_{\pi}^{2}(\mu)\big|_{\text{pert}} = \lim_{\vec{v} \to 0} \lim_{m \to \infty} \frac{3}{\vec{v}^{2}} \frac{\int_{0}^{\mu} d\omega \, \omega^{2} \, W(\omega, \vec{v})}{\int_{0}^{\mu} d\omega \, W(\omega, \vec{v})}$$

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• We have to calculate the imaginary part of the forward scattering amplitudes:



$$W(\omega, \vec{v}) = W_{\rm el}(\vec{v})\delta(\omega) + \frac{\vec{v}^2}{\omega}W_{\rm real}(\omega)\theta(\omega) + \mathcal{O}(v^4, \frac{\omega}{m_b})$$

Virtual corrections: only contribute to the denominator

- given by the static limit of the massive form factor
 [Ravidran, Neerven, PLB 445 (1998),..., Ablinger et al. PRD 97 (2018), ...]
- real corrections: only contribute to the numerator
 - needs to be computed in the appropriate limit

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• We can translate to a kovariant expansion:

$$egin{aligned} y&=m_b^2-s=m_b\omega(2+v^2)+\mathcal{O}(\omega^2,v^4)\ q^2&=-m_bv^2(m_b-\omega)+\mathcal{O}(\omega^2,v^4)\ 2p\cdot q&=y-q^2 \end{aligned}$$

- We realize the threshold expansion via expansion by regions: [Beneke, Smirnov (Nucl. Phys. B (1998))]
 - the loop momenta can scale: hard (h) $k_i \sim m_b$ or ultrasoft (u) $k_i \sim y/m_b$
 - not all *k_i* can be hard, otherwise no imaginary part is produced.
 - we checked that we reproduce all regions with these scalings with Asy.m [Pak, Smirnov (Eur. Phys. J. C (2011))]
- expansion in \vec{v} reduces to a Taylor expansion in q.



Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to two-loop on-shell master integrals.
 [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators

hard region ($k_1 \sim m_b$):

$$\int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][(p+q+k_1)^2 - m_b^2][(p+k_1)^2 - m_b^2]}$$

= $\int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][k_1^2 + 2(p+q) \cdot k_1 - y][k_1^2 + 2p \cdot k_1]}$
 $\rightarrow \int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][k_1^2 + 2p \cdot k_1]^2} + \dots$

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 [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators

ultrasoft region ($k_1 \sim y/m_b$):

$$\int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][(p+q+k_1)^2 - m_b^2][(p+k_1)^2 - m_b^2]}$$

= $\int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][k_1^2 + 2(p+q) \cdot k_1 - y][k_1^2 + 2p \cdot k_1]}$
 $\rightarrow \int \frac{d^d k_1}{(2\pi)^d} \frac{1}{[k_1^2][2p \cdot k_1 - y][2p \cdot k_1]} + \dots$

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Calculation of the Master Integrals – ultrasoft regions



Three types of propagators left:

- massless propagators:
 k_i² (dotted lines)
- eikonal propagators:
 2k_i.p (double lines)
- 'massive' eikonal propagators: $2k_{i}.p y$ (solid lines)



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Strategy:

direct integration of Feynman parameters in terms of Γ-functions

$$= \prod_{i=1}^{3} \int \frac{\mathrm{d}k_i}{(2\pi)^d} \frac{1}{[-k_1^2][-(k_1+k_3)^2][-(k_2+k_3)^2][-2p.k_2][-2p.k_1+y]} }{ = y^{3d-8}m^{6-3d}\frac{\Gamma^3(d/2-1)\Gamma(5-2d)\Gamma(3d-7)\Gamma(8-3d)}{\Gamma(d-2)} }{ = y^{4-6\epsilon}m^{6\epsilon-6}S_{\epsilon}^3 \left\{ -\frac{1}{144\epsilon^2} - \frac{7}{108\epsilon} - \frac{127}{324} - \frac{29\pi^2}{576} - \epsilon\left(\frac{478}{243} + \frac{203}{432}\pi^2 - \frac{23}{144}\zeta_3\right) + \dots \right\}$$

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Strategy:

- We can find an up to 2-fold Mellin-Barnes representation.
- Two approaches for the solution:
 - I High precision numerical integration together with PSLQ.
 - Analytic summation of residue sums with the help of Sigma, EvaluateMultiSums and HarmonicSums [Schneider (2007-); Ablinger et al (2011-)].

$$= \prod_{i=1}^{3} \int \frac{dk_i}{(2\pi)^d} \frac{1}{[-k_1^2][-k_3^2][-(k_1-k_2)^2][-(k_2-k_3)^2][-2p.k_1][-2p.k_3][-2p.k_2+y]}$$

$$= y^{3d-11} m^{8-3d} \frac{\Gamma(11-3d)\Gamma(3d-10)}{\Gamma^2(d-3)} \left(\frac{1}{2\pi i}\right)^2 \int_{-i\infty}^{+i\infty} dw_1 \int_{-i\infty}^{+i\infty} dw_2 \dots$$

$$= y^{1-6\epsilon} m^{6\epsilon-4} S_{\epsilon}^3 \left\{ -\frac{\zeta_3}{\epsilon} - \frac{3}{5}\zeta_2^2 - 8\zeta_3 + \epsilon \left(\frac{9}{2}\zeta_2\zeta_3 - \frac{24}{5}\zeta_2^2 - 52\zeta_3 - 57\zeta_5\right) + \mathcal{O}(\epsilon^2) \right\}$$

$$= v^{1-6\epsilon} m^{6\epsilon-4} S_{\epsilon}^3 \left\{ -\frac{\zeta_3}{\epsilon} - \frac{3}{5}\zeta_2^2 - 8\zeta_3 + \epsilon \left(\frac{9}{2}\zeta_2\zeta_3 - \frac{24}{5}\zeta_2^2 - 52\zeta_3 - 57\zeta_5\right) + \mathcal{O}(\epsilon^2) \right\}$$

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Strategy:

- We find a 3- or 4-fold Mellin-Barnes representation.
- Approache for the solution:
 - Introduce a new scale $x \neq y$ into some of the massive eikonal propagators.
 - Derive differential equations and solve with the help of Sigma, OreSys and HarmonicSums [Schneider (2007-); Ablinger et al (2011-)].
 - Fix boundary for $x \to 0$.



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Strategy:

• We find a 3- or 4-fold Mellin-Barnes representation.

$$= \prod_{i=1}^{3} \int \frac{\mathrm{d}k_{i}}{(2\pi)^{d}} \frac{1}{[-k_{1}^{2}][-k_{3}^{2}][-(k_{1}-k_{2})^{2}][-(k_{2}-k_{3})^{2}][-2\rho.k_{1}+y][-2\rho.k_{3}+y][-2\rho.k_{2}+y]}{= y^{1-6\epsilon} m^{6\epsilon-2} \left(\frac{1}{2\pi i}\right)^{4} \int_{-i\infty}^{+i\infty} dw_{1} \int_{-i\infty}^{+i\infty} dw_{2} \int_{-i\infty}^{+i\infty} dw_{3} \int_{-i\infty}^{+i\infty} dw_{4} \dots \right.$$

$$= y^{1-6\epsilon} m^{6\epsilon-2} \left\{ -\frac{\zeta_{2}}{3\epsilon^{2}} + \frac{1}{\epsilon} \left(\frac{\zeta_{3}}{3} - \frac{8\zeta_{2}}{3}\right) - \frac{52\zeta_{2}}{3} + \frac{8\zeta_{3}}{3} - \frac{49\zeta_{2}^{2}}{30} + \epsilon \left(-\frac{320}{3}\zeta_{2} + \frac{52}{3}\zeta_{3} - \frac{196}{15}\zeta_{2}^{2} - \frac{11}{6}\zeta_{2}\zeta_{3} - \frac{83}{3}\zeta_{5}\right) + \mathcal{O}(\epsilon^{2}) \right\}$$

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Charm mass dependence





Previously no charm mass effects for the kinetic mass had been known.

- We assume $|y| \ll m_c^2, m_b^2$, i.e. no cuts through a charm loop.
- The bare result has a non-trivial dependence on m_c .
- After renormalization only decoupling effects are left.
- \Rightarrow The kinetic mass conversion has no explicit m_c dependence if parametrized in terms of $\alpha_s^{(3)}$.

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Results



$$\begin{split} \frac{m^{\text{kin}}}{m^{\text{OS}}} &= 1 - \frac{\alpha_s^{(n_l)}}{\pi} C_F \left(\frac{4}{3} \frac{\mu}{m^{\text{OS}}} + \frac{1}{2} \frac{\mu^2}{(m^{\text{OS}})^2}\right) + \left(\frac{\alpha_s^{(n_l)}}{\pi}\right)^2 C_F \left\{\frac{\mu}{m^{\text{OS}}} \left[C_A \left(-\frac{215}{27} + \frac{2\pi^2}{9} + \frac{22}{9}l_\mu\right) + n_l T_F \left(\frac{64}{27} - \frac{8}{9}l_\mu\right)\right]\right] \\ &+ \frac{\mu^2}{(m^{\text{OS}})^2} \left[C_A \left(-\frac{91}{36} + \frac{\pi^2}{12} + \frac{11}{12}l_\mu\right) + n_l T_F \left(\frac{13}{18} - \frac{1}{3}l_\mu\right)\right]\right\} + \left(\frac{\alpha_s^{(n_l)}}{\pi}\right)^3 C_F \left\{\frac{\mu}{m^{\text{OS}}} \left[C_A^2 \left(-\frac{130867}{1944} + \frac{511\pi^2}{162} + \frac{19\zeta_3}{2} - \frac{\pi^4}{18} + \left(\frac{2518}{81} - \frac{22\pi^2}{27}\right)l_\mu - \frac{121}{27}l_\mu^2\right) + C_A n_l T_F \left(\frac{19453}{486} - \frac{104\pi^2}{81} - 2\zeta_3 + \left(-\frac{1654}{81} + \frac{8\pi^2}{27}\right)l_\mu + \frac{88}{27}l_\mu^2\right) + C_F n_l T_F \left(\frac{11}{4} - \frac{4\zeta_3}{3} - \frac{2}{3}l_\mu\right) + n_l^2 T_F^2 \left(-\frac{1292}{243} + \frac{8\pi^2}{81} + \frac{256}{81}l_\mu - \frac{16}{27}l_\mu^2\right)\right] \\ &+ \frac{\mu^2}{(m^{\text{OS}})^2} \left[C_A^2 \left(-\frac{96295}{5184} + \frac{445\pi^2}{432} + \frac{57\zeta_3}{16} - \frac{\pi^4}{48} + \left(\frac{2155}{216} - \frac{11\pi^2}{36}\right)l_\mu - \frac{121}{72}l_\mu^2\right) + C_A n_l T_F \left(\frac{13699}{1296} - \frac{23\pi^2}{54} - \frac{3\zeta_3}{4} + \left(-\frac{695}{108} + \frac{\pi^2}{9}\right)l_\mu + \frac{11}{9}l_\mu^2\right) + C_F n_l T_F \left(\frac{29}{32} - \frac{\zeta_3}{2} - \frac{1}{4}l_\mu\right) + n_l^2 T_F^2 \left(-\frac{209}{162} + \frac{\pi^2}{27} + \frac{26}{27}l_\mu - \frac{2}{9}l_\mu^2\right)\right]\right\}, (4) \end{split}$$

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Results



Using as inputs $\overline{m}_b(\overline{m}_b) = 4163 \text{ MeV}$, $\alpha_s^{(5)}(M_Z) = 0.1179$:

$$\begin{array}{ll} m_c = 0: \\ n_l = 3: \\ n_l = 4: \end{array} & \begin{array}{ll} m_b^{\rm kin}(1 \, {\rm GeV}) = (4163 + 248 + 81 + 30) \, {\rm MeV} = 4521(15) \, {\rm MeV} \\ m_b^{\rm kin}(1 \, {\rm GeV}) = (4163 + 259 + 77 + 25) \, {\rm MeV} = 4523(12) \, {\rm MeV} \end{array}$$

To be compared with:

• scheme conversion uncertainty at two loops: $\delta m_b^{\rm kin} =$ 30 MeV [Gambino, JHEP 09 (2011)]

• m_b from $b \rightarrow c\ell\nu$ global fit: $m_b^{\rm kin}(1\,{
m GeV}) = 4554\pm18\,{
m MeV}$ [HFLAV, EPJC 81 (2021)]

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Results



Using as inputs
$$\overline{m}_b(\overline{m}_b) = 4163 \text{ MeV}, \ \alpha_s^{(5)}(M_Z) = 0.1179$$
:

$m_{\rm c} = 0$: $n_l = 3$: $m_b^{\rm kin}(1\,{\rm GeV}) = (4163 + 248 + 81 + 30)\,{\rm MeV} = 4521(15)\,{\rm MeV}$ $m_{\rm b}^{\rm kin}(1\,{\rm GeV}) = (4163 + 259 + 77 + 25)\,{\rm MeV} = 4523(12)\,{\rm MeV}$ $n_l = 4$: $m_c \neq 0$: $m_b^{\rm kin}(1\,{\rm GeV}) = (4163 + 248 + 80 + 30)\,{\rm MeV} = 4520(15)\,{\rm MeV}$ $n_{l} = 3$: $n_l = 4$: $m_{\rm b}^{\rm kin}(1\,{\rm GeV}) = (4163 + 259 + 78 + 26)\,{\rm MeV} = 4526(12)\,{\rm MeV}$

To be compared with:

- scheme conversion uncertainty at two loops: $\delta m_b^{\rm kin} = 30 \,{\rm MeV}$ [Gambino, JHEP 09 (2011)]
- m_b from $b \rightarrow c \ell \nu$ global fit: $m_b^{\rm kin}$ (1 GeV) = 4554 \pm 18 MeV [HFLAV, EPJC 81 (2021)]

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Method of Calculation



- We calculate the inclusive decay rate to third order via the optical theorem, i.e. we consider the imaginary part of 5-loop forward scattering diagrams.
- We consider massless leptons, i.e. we have two dimensionful scales, the bottom mass m_b and the charm mass m_c.
- Analytical dependence on charm and bottom mass seems out of reach:
 - \Rightarrow consider expansion in mass difference



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The Heavy-Daughter Expansion



- Perform the expansion in the limit $m_c \sim m_b$: $\delta = 1 \rho = 1 \frac{m_c}{m_b} \ll 1$
- Limit has been shown to converge well down to $m_c/m_b \rightarrow 0$ at 2-loop order. [Czarnecki, Dowling, Piclum (Phys. Rev. D 78 (2008))]



Asymptotic Expansion



The expansion in $\delta = 1 - m_c/m_b$ is very similar to the threshold expansion for the kinetic mass:

- We use the method of regions to perform the expansion. [Beneke, Smirnov (Nucl. Phys. B (1998))]
- Loop momenta can either scale hard k_i ~ m_b or ultra-soft k_i ~ δm_b (regions have been cross-checked with Asy). [Pak, Smirnov (Eur. Phys. J. C (2011))]
- The momentum of the electron-neutrino loop can be integrated trivially.
- The properties of the asymptotic expansion allow to factorize the leptonic system completely.
- \Rightarrow We can reduce our 5-loop to 3-loop integrals with integer powers without any integration-by-parts.

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Asymptotic Expansion – Example



Look at the 1-loop integral (we already integrated out the electron-neutrino loop):

$$\sim \int \frac{\mathrm{d}q \mathrm{d}k}{[q^2]^{\alpha}[(p+q)^2 - m_c^2]^2[k^2][(p+q+k)^2 - m_c^2]}$$

• case 1: q has to be ultra-soft, k is hard $(q \sim \delta m_b, k \sim m_b)$;

$$\rightarrow \int \frac{\mathrm{d}q}{[q^2]^{\alpha} [2p \cdot q + 2m_b^2 \delta]^2} \times \int \frac{\mathrm{d}k}{[k^2][(k+p)^2 - m_b^2]}$$

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Asymptotic Expansion – Example

q



Look at the 1-loop integral (we already integrated out the electron-neutrino loop):

$$\sim \int \frac{\mathrm{d}q \mathrm{d}k}{[q^2]^{\alpha}[(p+q)^2 - m_c^2]^2[k^2][(p+q+k)^2 - m_c^2]}$$

• case 2: q and k are ultra-soft (q, $k \sim \delta m_b$);

$$\rightarrow \int \frac{\mathrm{d}q \mathrm{d}k}{[q^2]^{\alpha} [2p \cdot q + 2m_b^2 \delta]^2 [k^2] [2p \cdot k + \underbrace{2p \cdot q + 2m_b^2 \delta}_{\text{fixed combination}}]$$

$$= \int \frac{\mathrm{d}q}{[q^2]^{\alpha} [2p \cdot q + 2m_b^2 \delta]^{\beta}} \times \int \frac{\mathrm{d}k}{[k^2] [2p \cdot k + 1]}$$

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Details on the Calculation



We can always perform the integrations over the electron-neutrino loop and lepton momenta analytically via 1-loop tensor reduction. The remaining loop integration have the following scalings:

	scaling	n. regions
$\mathcal{O}(\alpha_s)$	h, u	2
$\mathcal{O}(\alpha_s^2)$	hh, hu, uu	4
$\mathcal{O}(\alpha_s^3)$	hhh, huu, hhu, uuu	8

- In case a single region with either hard or ultra-soft scaling remains we can also integrate it out analytically.
- The remaining two- or three-loop integrals have integer powers of the propagators and can be reduced to master integrals via IBP reduction.

 $\begin{array}{c} \text{Introduction} & \text{The Kinetic Mass} \\ \text{ooo} & \text{oooooooooooo} \\ \text{Kay Schönwald} - \mathcal{O}(\alpha_s^3) \text{ calculations for the inclusive determination of } |V_{cb}| \\ \end{array}$

Details on the calculation



- We have to consider 1450 five-loop diagrams.
- Several subtleties with FORM:
 - Major obstacle is to keep the size of FORM expressions as low as possible.
 - Efficient expansion of denominators and memory management.
 - Automated partial fractioning and sector/families mapping. [LIMIT, F. Herren, PhD Thesis, KIT, 2020]
- Intermediate FORM expressions of O(100) GB.
- About 25M three-loop integrals with positive and negative indices up to 12 had to be reduce. We used (a private version of) FIRE together with LiteRed (also standalone).

Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to three loop on-shell master integrals.
 [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with (massive) eikonal propagators

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Details on the calculation



- We have to consider 1450 five-loop diagrams.
- Several subtleties with FORM:
 - Major obstacle is to keep the size of FORM expressions as low as possible.
 - Efficient expansion of denominators and memory management.
 - Automated partial fractioning and sector/families mapping. [LIMIT, F. Herren, PhD Thesis, KIT, 2020]
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Renormalization



For the renormalization of the decay width we need

- the wave function renormalization constant Z₂
- the mass renormalization constant Z_m

with two massive quarks in the expansion $m_c \sim m_b$ up to $O(\alpha_s^3)$.

• Previously they were only known in the expansion $x = m_c/m_b \sim 0$ and numerically for larger values of m_c [Bekavac, Grozin, Seidel, Steinhauser, JHEP 10 (2007)].



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Definition



• The masses and the wave function can be renormalized multiplicatively:

$$m^0 = Z_m^{OS} m_{OS}, \qquad \qquad \psi^0 = \sqrt{Z_2^{OS} \psi^{OS}}$$

• The renormalization constants can be calculated from the quark self energy:

$$\Sigma(q^2,m)=m\Sigma_1(q^2,m)+(\not q-m)\Sigma_2(q^2,m).$$



Computation Strategy

- Generate graphs with qgraf.
- Generate amplitudes with q2e/exp.
- Reduce to master integrals with Crusher and LiteRed.
- Derive a closed system of differential equations in the variable $x = m_c/m_b$ for the master integrals and solve.



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Computation Strategy



To solve the coupled system of differential equations we use the algorithm presented in [Ablinger, Blümlein, Marguard, Rana, Schneider, NPB 939 (2019)] :

- Uncouple coupled systems using OreSys to a single differential equation of higher order.
- Solve the higher order differential equations by factorizing the differential operator (as implemented in HarmonicSums).
- Fix the boundary constants from the known values at x = 0 and/or x = 1.

We find 4 master integrals which cannot be expressed in terms of harmonic

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- Solve the higher order differential equations by factorizing the differential operator (as implemented in HarmonicSums).
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 We find 4 master integrals which cannot be expressed in terms of harmonic polylogarithms.

Computation Strategy



 M_{23}

For example we find the homogenous differential equation:

$$M_{22}^{\prime\prime} = rac{1-4x^2}{x(1-x^2)}M_{22}^{\prime} + rac{4}{1-x^2}M_{22} \; .$$



$$M_{22}^{(1)} = x^2(4 - x^2) ,$$

$$M_{22}^{(2)} = (2 - 3x^2 + x^4)\sqrt{1 - x^2} + x^2(4 - x^2) \left[1 + G\left(\left\{\frac{\sqrt{1 - \tau^2}}{\tau}\right\}, x\right)\right] .$$

$$G(\{f(\tau), \vec{g}(\tau)\}, x) = \int_{0}^{x} d\tau_{1} f(\tau_{1}) G(\{\vec{g}(\tau)\}, \tau_{1})$$

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Computation Strategy



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$$\begin{split} M_{22}^{(1)} &= x^2 (4 - x^2) , \\ M_{22}^{(2)} &= (2 - 3x^2 + x^4) \sqrt{1 - x^2} + x^2 (4 - x^2) \left[1 + G\left(\left\{ \frac{\sqrt{1 - \tau^2}}{\tau} \right\}, x \right) \right] . \end{split}$$

$$G({f(\tau), \vec{g}(\tau)}, x) = \int_{0}^{x} d\tau_{1} f(\tau_{1}) G({\{\vec{g}(\tau)\}, \tau_{1}})$$

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Results - Analytic Structure



• For the full solution we need the letters:

$$\frac{1}{1-\tau}, \quad \frac{1}{\tau}, \quad \frac{1}{1+\tau}, \quad \sqrt{1-\tau^2}, \quad \frac{\sqrt{1-\tau^2}}{\tau}$$

• The last two letters can be rationalized with:

$$x=\frac{2y}{1+y^2},$$

then we have cyclotomic harmonic polylogarithms [Ablinger, Blümlein, Schneider, JMP 52 (2011)] :

$$\frac{1}{1-\tau}, \quad \frac{1}{\tau}, \quad \frac{1}{1+\tau}, \quad \frac{1}{1+\tau^2}, \quad \frac{\tau}{1+\tau^2}$$

• The constants at x = y = 1 are known (at least up to weight 5).

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Results - Series Expansions

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Results - Decay Width



$$\begin{split} \mathsf{F}(b \to c\ell\nu) &= \mathsf{F}_0 \left[X_0 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi} \right)^n X_n \right] \\ X_3 &= \delta^5 \left(\frac{266929}{810} - \frac{5248a_4}{27} + \frac{2186\pi^2 \zeta_3}{45} - \frac{4094\zeta_3}{45} - \frac{1544\zeta_5}{9} - \frac{656l_2^4}{81} + \frac{1336}{405}\pi^2 l_2^2 \right. \\ &+ \frac{44888\pi^2 l_2}{125} - \frac{9944\pi^4}{2025} - \frac{608201\pi^2}{2420} \right) + \delta^6 \left(-\frac{284701}{540} + \frac{2624a_4}{9} - \frac{1093\pi^2 \zeta_3}{15} \right) \end{split}$$

$$+ \frac{135}{3} + \frac{2025}{3} + \frac{2430}{27} + \frac{10}{35} \left(\frac{540}{9} + \frac{9}{15} + \frac{391\zeta_3}{3} + \frac{772\zeta_5}{3} + \frac{328l_2^4}{27} - \frac{668}{135} \pi^2 l_2^2 - \frac{1484\pi^2 l_2}{3} + \frac{4972\pi^4}{675} + \frac{591641\pi^2}{1620} \right) + \mathcal{O}(\delta^7 \ln^2(\delta)) ,$$

with
$$I_2 = \ln(2)$$
, $a_4 = \text{Li}_4(1/2)$, $\zeta_i = \sum_{j=1}^{\infty} j^{-i}$ and $\mu_s = m_b$.

• We have calculated the expansion up to δ^{12} (for general color factors).

• A subset of color factors has been independently been computed up to δ^9 .

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Results - Decay Width



$$\begin{split} & \Gamma(b \to c\ell\nu) = \Gamma_0 \left[X_0 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi} \right)^n X_n \right] \\ & X_3 = \delta^5 \left(\frac{266929}{810} - \frac{5248a_4}{27} + \frac{2186\pi^2\zeta_3}{45} - \frac{4094\zeta_3}{45} - \frac{1544\zeta_5}{9} - \frac{656l_2^4}{81} + \frac{1336}{405}\pi^2 l_2^2 \right. \\ & + \frac{44888\pi^2 l_2}{135} - \frac{9944\pi^4}{2025} - \frac{608201\pi^2}{2430} \right) + \delta^6 \left(-\frac{284701}{540} + \frac{2624a_4}{9} - \frac{1093\pi^2\zeta_3}{15} \right. \\ & + \frac{391\zeta_3}{3} + \frac{772\zeta_5}{3} + \frac{328l_2^4}{27} - \frac{668}{135}\pi^2 l_2^2 - \frac{1484\pi^2 l_2}{3} + \frac{4972\pi^4}{675} + \frac{591641\pi^2}{1620} \right) + \mathcal{O}(\delta^7 \ln^2(\delta)) \,, \end{split}$$

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[Czakon, Czarnecki, Dowling (2021)]

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Convergence – Quark Decays



• We see a good convergence at the physical point of $\rho = m_c/m_b \approx 0.28$.

We find:

 $X_3(
ho=0.28)=-68.4\pm0.3$

- We use the difference of the last two expansion terms to estimate the uncertainty.
- For $\rho \to 0$ we can extract values for $b \to u\ell\nu$:

$$\begin{array}{c} 0 \\ -50 \\ -100 \\ X \\ -150 \\ -200 \\ -200 \\ -250 \\ -300 \\ 0.0 \end{array} \begin{array}{c} 0.1 \\ 0.2 \\ 0.1 \\ 0.2 \\ 0.3 \\ 0.4 \\ 0.5 \end{array}$$

$$X_3^u = -202 \pm 20$$

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Different Renormalization Schemes



The total decay rate of quarks expressed in terms of on-shell masses converges poorly:

$$\Gamma_{
m sl} \sim 1 - 1.72 rac{lpha_s(m_b)}{\pi} - 13.1 \left(rac{lpha_s(m_b)}{\pi}
ight)^2 - 163 \left(rac{lpha_s(m_b)}{\pi}
ight)^3$$

• Also the $\overline{\mathrm{MS}}$ scheme usually behaves poorly, since the scale has to be chosen rather low.

- Different threshold masses like the PS [Beneke (1998)], 1S [Hoang, Ligeti, Manohar (1998)] or kinetic mass [Bigi, Shifman, Uraltsev, Vainshtein (1996)] have been proposed to improve the convergence.
- We see a much better behavior in the convergence for the schemes used for the global fits of inclusive quantities.
- E.g. for the kinetic mass:

$$m_b^{
m kin}, m_c^{
m kin}: \Gamma(b
ightarrow c \ell
u) / \Gamma_0 = 0.633 \left(1 - 0.066 - 0.018 - 0.007
ight) pprox 0.575$$

 $m_b^{
m kin}, \overline{m}_c(3~{
m GeV}):$ $\Gamma(b
ightarrow c \ell
u) / \Gamma_0 = 0.700 \left(1 - 0.116 - 0.035 - 0.010\right) \approx 0.587$

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Different Renormalization Schemes



BLM and non-BLM part

$$\Gamma(b \to c\ell\nu) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} X_0 \begin{bmatrix} 1 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi}\right)^n Y_n \end{bmatrix} \qquad \qquad Y_2 = \beta_0 Y_2^{\beta_0} + Y_2^{\text{rem}} \\ Y_3 = \beta_0^2 Y_3^{\beta_0} + Y_3^{\text{rem}} \end{bmatrix}$$

	<i>Y</i> ₁	Y_2^{rem}	$eta_{0} Y_{2}^{eta_{0}}$	$Y_3^{ m rem}$	$eta_0^2 Y_3^{eta_0^2}$
m_b^{OS}, m_c^{OS}	-1.72	3.08	-16.17	48.8	-212.1
$m_b^{ m kin}, m_c^{ m kin}$	-0.94	0.33	-4.08	-5.4	-15.4
$m_b^{ m kin}, \overline{m}_c$ (3 GeV)	-1.67	-3.39	-3.85	-97.7	69.1
$m_b^{ m kin}, \overline{m}_c$ (2 GeV)	-1.25	-1.21	-2.43	-68.8	67.9
$\overline{m}_b(\overline{m}_b), \overline{m}_c(3 \text{ GeV})$	3.07	-21.81	35.17	-56.7	119.4
$\mathit{m^{\mathrm{PS}}_b}, \overline{\mathit{m}}_{c}(2~GeV)$	-0.47	-6.10	-2.31	-93.1	-7.19
$m_b^{ m 1S}, \overline{m}_c({ m 2~GeV})$	-3.59	-0.98	-19.39	-39.83	-80.22
$\mathit{m^{\mathrm{1S}}_{b}}, \mathit{m_{c}}$ via HQET	-1.38	0.73	-7.05	5.04	-38.09

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Different Renormalization Schemes



BLM and non-BLM part

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Different Renormalization Schemes – kinetic mass





Different Renormalization Schemes – 1S mass





The Updated Fit



- experimental moments from 2014 [Belle,Babar,CDF,CLEO,DELPHI]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\Gamma(B \to X_c \ell \bar{\nu})$ [Fael,Schönwald,Steinhauser (2020)]
- $\mathcal{O}(\alpha_s^3)$ corrections to $\overline{m}_b m_b^{\rm kin}$ relation [Fael,Schönwald,Steinhauser (2020)]
- $\overline{m}_c(3 \text{ GeV}) = 0.988(7) \text{ GeV}$, $\overline{m}_b(\overline{m}_b) = 4.198(12) \text{ GeV} \longrightarrow m_b^{\text{kin}} = 4.565(19) \text{ GeV}$ [FLAG (2019)]

$$|V_{cb}| = 42.16(30)_{
m th}(32)_{
m exp}(25)_{\Gamma} imes 10^{-3}$$

- $\Gamma(B \to X_c \ell \bar{\nu})_{\mathcal{O}(\alpha_s^3)}$: shift: $|V_{cb}|$ by +0.6% reduce uncertainty: $(50)_{\Gamma} \Rightarrow (25)_{\Gamma}$
- 1.2% uncertainty
- (32)_{exp} I improvements from Belle II

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The Updated Fit



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$$|\textit{V}_{\textit{cb}}| = 42.16(30)_{\rm th}(32)_{\rm exp}(25)_{\Gamma} \times 10^{-3}$$

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Moments of Differential Distributions



The method can be used to calculate inclusive moments of differential distributions.
 For example we can calculate-q² moments:



Moments of Differential Distributions



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 For example we can calculate-q² moments:



Convergence – Muon Decays



- Specifying the color factor to QED and setting $\rho = m_e/m_\mu \approx 0$ we get the 3-loop contributions to the muon decay.
- We find:

$$X_3^\mu = -15.3 \pm 2.3$$

This leads to the shift:

 $\Delta au_{\mu} = (-9 \pm 1) \cdot 10^{-8} \, \mu s$

The current experimental value reads:



 $au_{\mu} =$ (2.1969811 \pm 0.0000022) μs

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Conclusions and Outlook



Conclusions

- We have calculated the relation between the kinetic and on-shell mass up to $\mathcal{O}(\alpha_s^3)$.
- We have computed the α_s^3 corrections to the width of $b \to c \ell \nu$.
- We performed an expansion in the limit $1 m_c/m_b \ll 1$ and demonstrated its good convergence.
- The result is one of the few third order corrections involving two mass scales.
- The results have been used to improve the inclusive determination of $|V_{cb}$.
- The results are also relevant for $b
 ightarrow u \ell \nu$ and the muon decay.

Outlook

The method of calculation can be applied for the calculation of moments of the differential distributions.

 $\begin{array}{c} \text{Introduction} & \text{The Kinetic Mass} \\ \text{ooo} & \text{oooooooooooo} \\ \text{Kay Schönwald} - \mathcal{O}(\alpha_{s}^{3}) \text{ calculations for the inclusive determination of } | V_{ab} \\ \end{array}$

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