

# Calculating compact binary dynamics

using effective field theory and computer algebra

Jan Steinhoff

Max-Planck-Institute for Gravitational Physics (AEI), Potsdam-Golm, Germany



DESY Zeuthen, January 9th, 2019

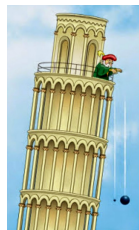
# Gravity: from Galileo and Newton to Einstein

Newton's theory of gravity:

- things fall down due to a gravitational force
- universality of free fall



London Science Museum



Einstein's theory of gravity

a.k.a. general relativity

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- free fall = force-free motion
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- space & time are curved

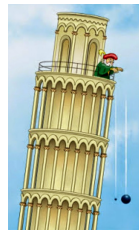
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International Space Station, nasa.gov

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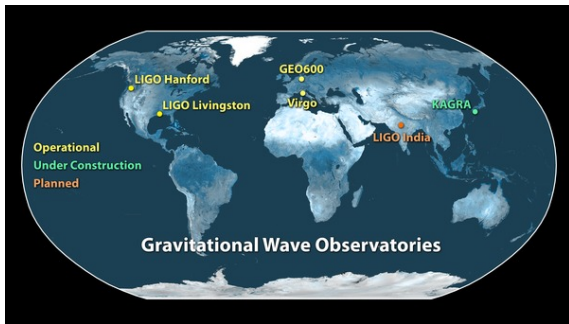


waves from a binary black hole  
[simulation with the Einstein Toolkit]

detector network:



Virgo, [www.ligo.caltech.edu/images](http://www.ligo.caltech.edu/images)



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# Detections of gravitational waves (GWs)

<https://www.ligo.caltech.edu/page/detection-companion-papers>

using LIGO:

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→ Nobel Prize in physics 2017

many more since then. . .

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→ (likely) a binary neutron star inspiral!

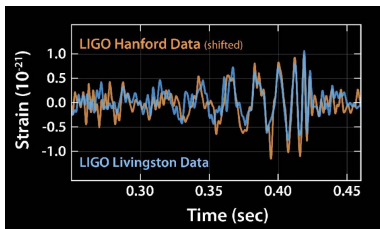
now: LIGO and Virgo

→ network of detectors still growing

→ accuracy/quantity will improve

e.g., drastically improved sky localization

→ need better predictions, too!



GW150914 (binary black hole), [ligo.caltech.edu](https://ligo.caltech.edu)

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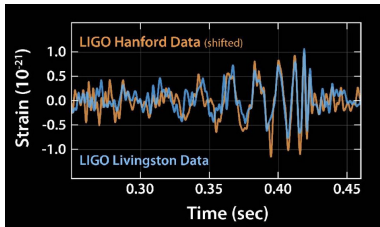
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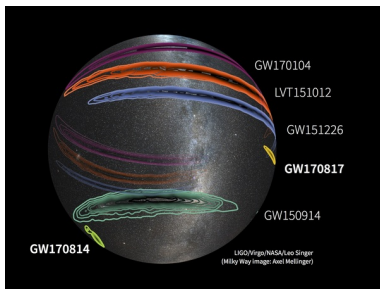
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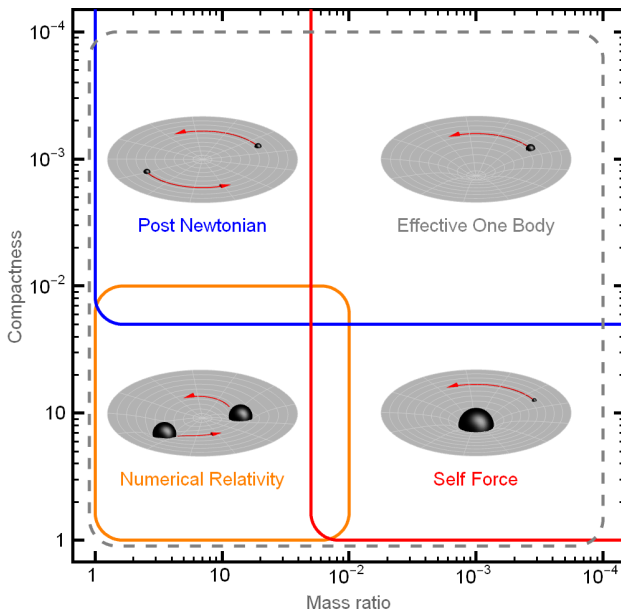
GW150914 (binary black hole), [ligo.caltech.edu](http://ligo.caltech.edu)



sky location of GW detections, [ligo.caltech.edu](http://ligo.caltech.edu)



# Method for predicting gravitational waves



wikipedia.org

# Effective field theory for post-Newtonian approximation

[Goldberger, Rothstein, PRD **73** (2006) 104029; Goldberger, arXiv:hep-ph/0701129]

for generic binaries in  
general relativity (GR):

- black holes (BHs)
- neutron stars (NSs)
- bold astronauts. . .

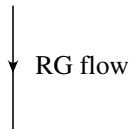
separation of scales:

- scale  $\mu$
- object size  $r_s$
- orbital size  $r$
- velocity  $v$   
 $\rightarrow$  frequency  $\sim \frac{v}{r}$

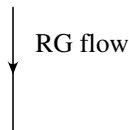
decomposition of the metric:

$$g_{\mu\nu} - \eta_{\mu\nu} = \left[ h_{\mu\nu}^s \sim \frac{1}{r_s} \right] + \left[ H_{\mu\nu} \sim \frac{1}{r} \right] + \left[ h_{\mu\nu} \sim \frac{v}{r} \right]$$

$$\mu = 1/r_s \frac{\text{BH, NS + GR}}{\text{pt. particle + GR}} \quad \text{match}$$



$$\mu = 1/r \frac{\text{bd. state}}{\text{mult. + rad.}} \quad \text{match}$$



$$\mu = v/r \text{ -----}$$

from [Goldberger, arXiv:hep-ph/0701129]

# Integrating out the orbital scale

action after integrating out  $h_{\mu\nu}^s$ : point-particles coupled to GR

$$S = -\frac{1}{16\pi G} \int d^4x \sqrt{\bar{g}} \bar{R} - \sum_{A=1,2} \int d\sigma_A \sqrt{\bar{g}_{\mu\nu} \frac{dx_A^\mu}{d\sigma_A} \frac{dx_A^\nu}{d\sigma_A}} (m_A + \dots)$$

$$\bar{g}_{\nu\nu} = \eta_{\mu\nu} + H_{\mu\nu} + h_{\mu\nu}$$

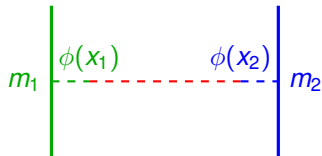
from now on: ignore radiation, drop  $h_{\mu\nu}$  contributions

add gauge fixing, post-Newtonian expansion:

$$S = -\frac{1}{16\pi G} \int dt d^3x \left[ 2\delta^{ij} \partial_i \phi \partial_j \phi + \dots \right] - \sum_A m_A \int dt \left[ 1 - \frac{\dot{\vec{x}}_A^2}{2} + \phi + \dots \right]$$
$$\bar{g}_{\mu\nu} dx^\mu dx^\nu = e^{2\phi} (dt - \mathbf{A}_i dx^i)^2 - e^{-2\phi} (\delta_{ij} + \sigma_{ij}) dx^i dx^j$$

integrate out  $\phi, \mathbf{A}_i, \sigma_{ij} \dots$

# Explicit example: the Newtonian limit

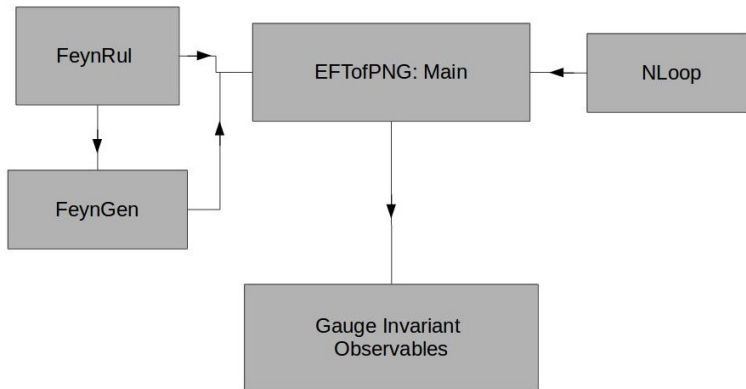


Newtonian potential as  
graviton exchange

$$\begin{aligned} &\approx \int dt_1 m_1 e^{i\vec{k} \cdot \vec{x}_1(t_1)} 4\pi G \delta(t_1 - t_2) \frac{d^3 k}{(2\pi)^3} \frac{1}{k^2} dt_2 m_2 e^{-i\vec{k} \cdot \vec{x}_2(t_2)} \\ &= \int dt \frac{G m_1 m_2}{r_{12}(t)} \quad \text{where } r_{12} = |\vec{x}_1 - \vec{x}_2| \end{aligned}$$

# Automation using computer algebra: EFTofPNG

M. Levi, JS, CQG **34** (2017) 244001, arXiv:1705.06309



## demo time

```
git clone -b talk-zeuthen20 https://github.com/jsteinhoff/pncbc-efctofpng.git
```

- wrap things into package environments, more abstractions
- need to improve substitution/representation of integrals
  - problem solved elsewhere (?)
  - delegate to other packages?
- optimal strategy to remove momentum-conserving  $\delta$ 's?
- more sophisticated way of collecting/sorting terms
  - keep subexpressions small
  - keep dim.-dependent prefactors together
- parallelize
  - should be straightforward, but memory is the limiting factor
- run with open source replacement for Mathematica?  
<https://mathics.github.io>